A Ground-Based 2-Micron DIAL System to Profile Tropospheric CO₂ and Aerosol Distributions for Atmospheric Studies

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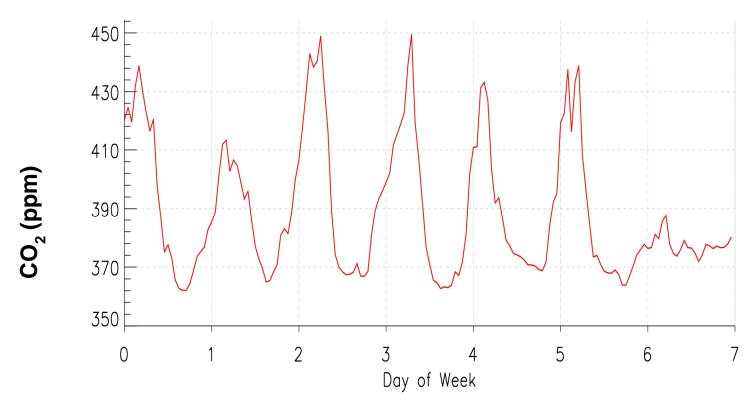
SPIE Asia-Pacific Remote Sensing, Goa, India23rd ILRC, NARA, Japan 13-17 November 2006

Acknowledge NASA Earth Science Technology Office (ESTO) funding

Objectives

- Technology development for profiling atmospheric CO₂
- •Coverage over 0-5 km altitude range with 0.5% precision
- •Deployable system for field studies, and as an anchor and validation tool for total column retrievals from the Orbiting Carbon Observatory (OCO)

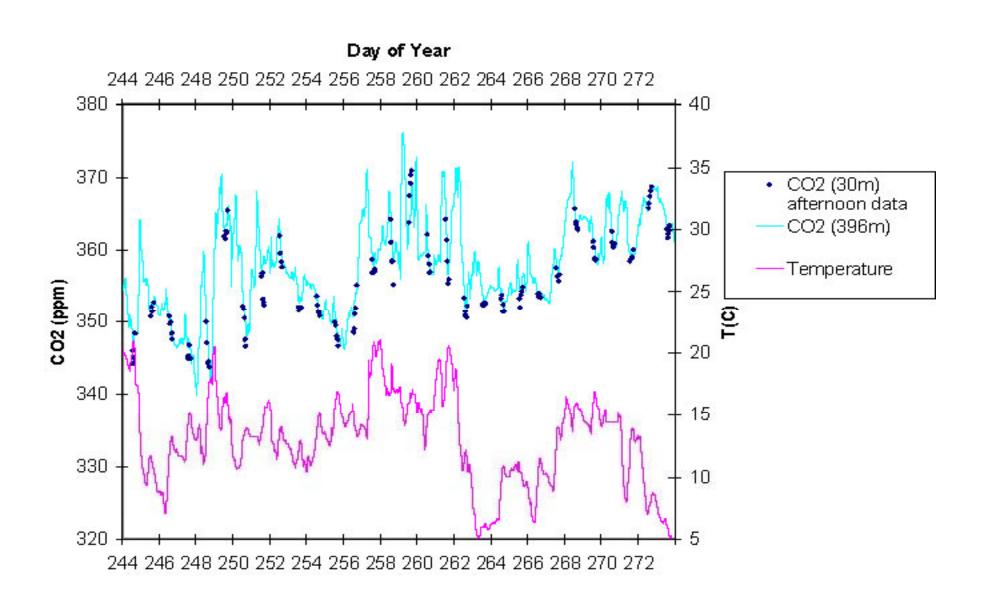
Need for Profiling of CO₂



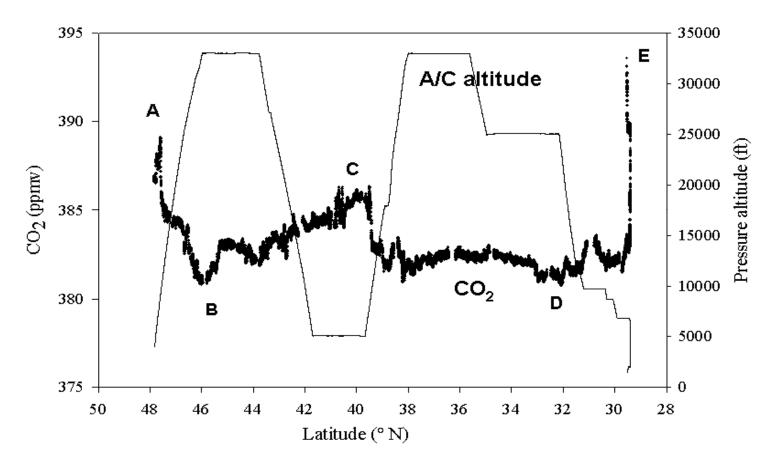
Diurnal variations of CO₂ using a NASA GSFC in situ LI-6262 sensor during the week of May 30, 2004. (Amelia Colarco, 2006)

Sources and sinks of CO₂ are located near the surface; and CO₂ is redistributed by passage of weather fronts, convection, dynamics, and transport

ABL CO₂ variations and their relation to synoptic weather

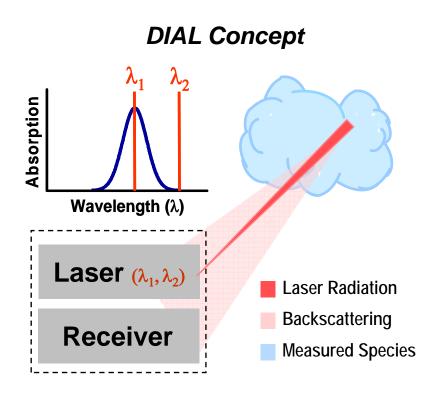


Advantage of CO₂ DIAL Profiling System



Variations in CO₂ observed along DC-8 flight track during INTEX-B using LI-6252 (Vay). Preliminary in situ data reveal A) BL to free trop transition, B) stratospheric air, C) biogenic and anthropogenic sources, D) tropical air, and E) free trop to BL transition.

Differential Absorption Lidar (DIAL) Technique and its Advantages



Advantages:

- High vertical and horizontal resolution
- •DIAL data permit direct inversion and absolute concentration measurements
- Simultaneous species and aerosol profiles, and cloud distributions
- Day and night coverage and no dependence on external radiation

DIAL technique for Atmospheric CO₂ measurements requires suitable absorption lines; narrow-band, tunable, and line locked lasers; high efficiency and low noise detectors

Features

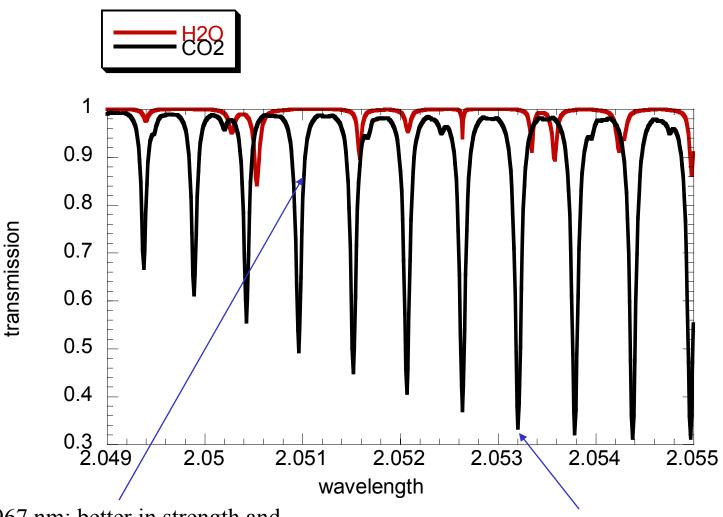
- •Excellent CO₂ absorption lines at 2.05 microns
- •State-of-the-art 2-mµ pulsed lasers
- •High QE, low noise, high gain detectors
- •Large collection area receiver

CO₂ absorption line selection and criteria

Select line that (i) allows optimum sampling of low troposphere, (ii) with low temperature sensitivity for differential absorption cross-section measurements, and (iii) has minimum interference from other species (in this case H₂O)

CO₂ lines in 2-micron region are very strong and require side-line operation. This requires line locking that will be off-set from line center.

Wavelength Selection



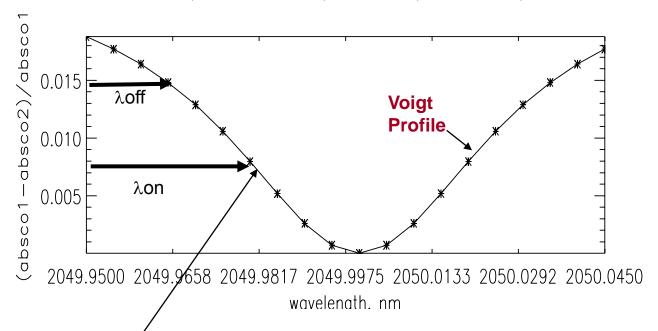
2050.967 nm: better in strength and temperature insensitivity. Moving to this line and operate on 'side-line'.

2053.204 nm: line described in proposal

Temperature insensitive CO₂ line selection

The $\Delta\sigma/\Delta T$ for mixing ratio measurements of the selected line R30 at 2050.967 nm*

CO2: 2050.0 nm, Is= 0.1741e-23, Iw= 0.070, eIs=362.788, Iwtd= 0.77



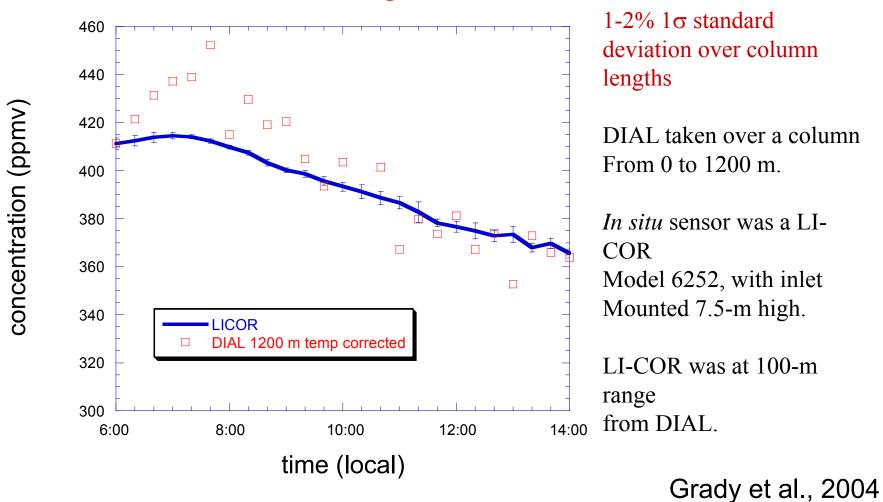
Sensitivity of absorption cross section σ to a change of 5 K temperature

Note: The $\Delta\sigma/\Delta T$ is very small ~0.14%/K, and even that is reduced by the off-line

^{*}A nominal wavelength was used in the temperature sensitivity calculations, and it has no impact on the results that are presented

Heritage: CO₂ concentration measured by DIAL Heterodyne detection using a 4" diameter telecope

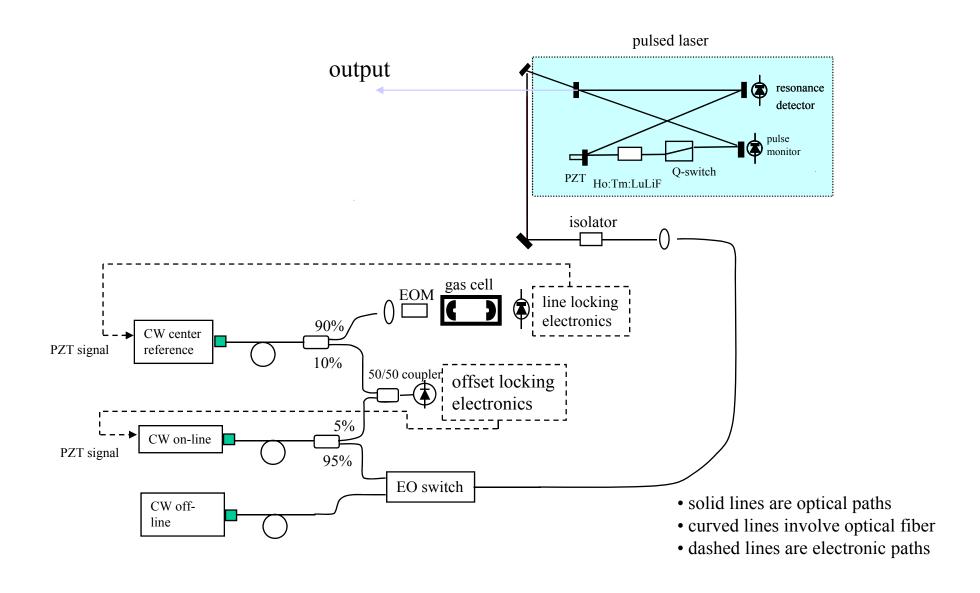




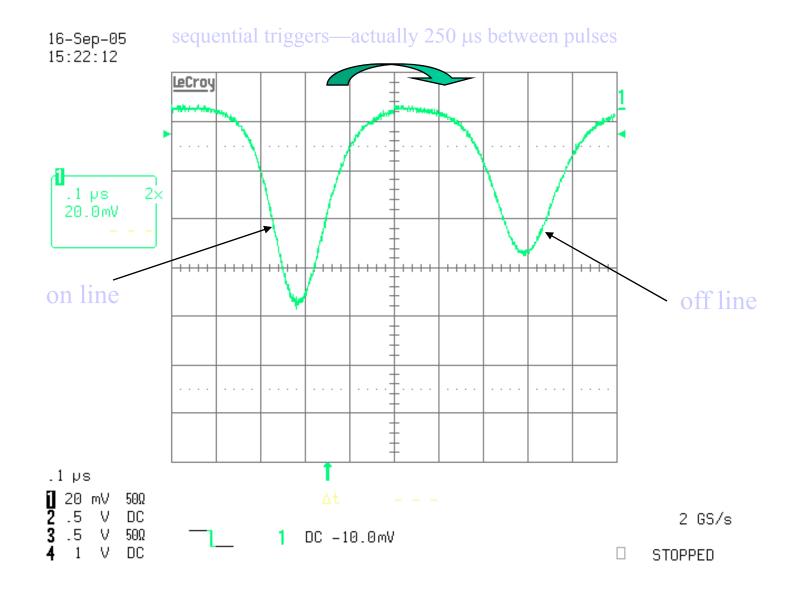
Laser Developments under earlier NASA Programs Including LRRP

- •Pulsed laser for range resolved profiling
- •Double pulsed operation to sample the same air mass by on- and off-laser pulses
- •Wavelength stability and spectrally narrow output
- •Line-locking with respect to a selected CO₂ line
- •Operation on a side of the line for optimum absorption cross-section selection

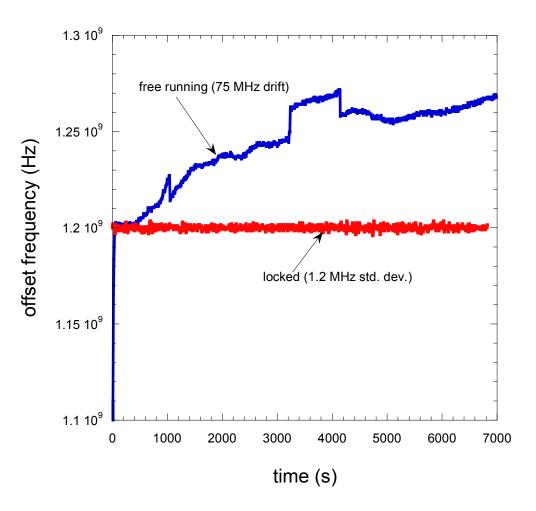
Transmitter System Architecture



Double-Pulse Wavelength-Switched Results

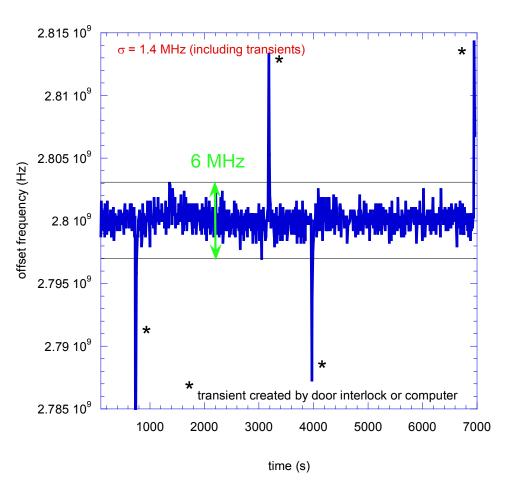


Side Line Offset Locking Characterization



- Electronic control holds an offset from center-line locked laser.
- Offset can be electronically programmed.
- Test here assesses quality of offset lock set for 1.2 GHz (37.3 pm).

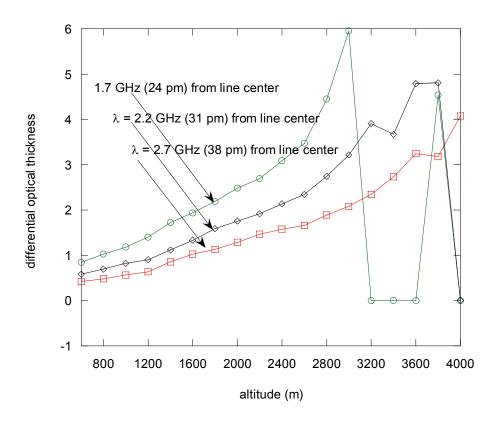
Side Line Offset Locking Characterization



- Electronic control holds an offset from center-line locked laser.
- Offset can be electronically programmed.
- Test here assesses quality of offset lock set for 2.8 GHz (37.3 pm).

Atmospheric test: changing side-line tuning, continued

diff. opt. thick. = $\ln (S_{off}/S_{on})$



- beam vertically pointing
- processed on 200-m range bins.
- •2000 pulse pairs.
- three sets of data
- optical depth changes with wavelength, demonstrating the usefulness of this technique.

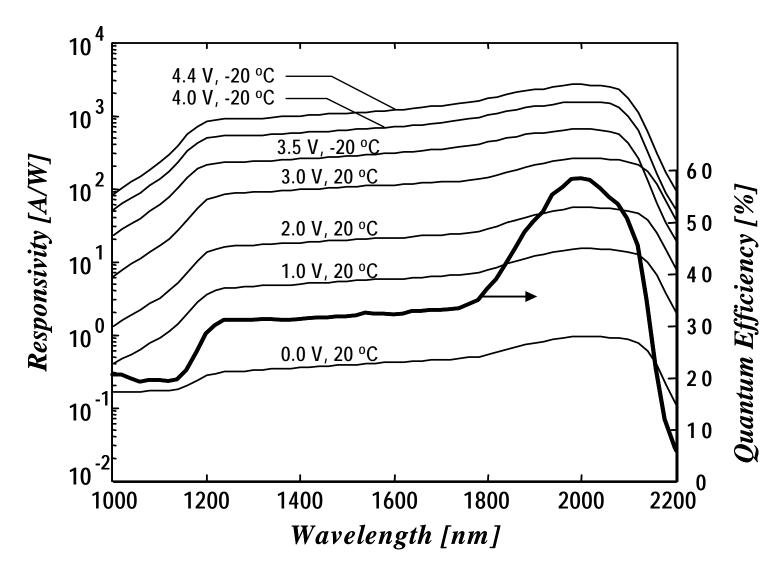
CO₂ DIAL (IIP) Goals/Requirement for detector

Advanced Detectors development needed that exceeds the performances of the commercially available detectors

- ➤ Responsivity @ 2.00-µm ≥ 50 A/W
- ➤ Operating temperature, T ~ 20 C
- \blacktriangleright Low noise (NEP \le 2 to 5x10 ⁻¹⁴ W/ \sqrt{Hz}) for far field
- ➤ Quantum Efficiency @ 2.00-µm ≥ 50%
- ➤ Bandwidth ≥ 5 MHz
- Collecting area diameter: 200 1000 μm

New AlGaAsSb/InGaAsSb phototransistors acquired and characterized at SED under LRRP—these showed high sensitivity at 2 micron wavelength.

AlGaAsSb/InGaAsSb Phototransistor With High Sensitivity at 2 µm



This detector has small (200 µ diameter) collection area

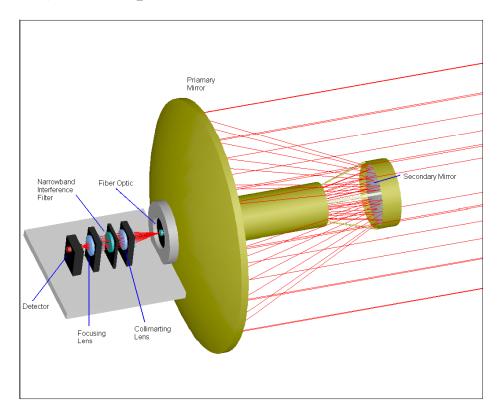
Design considerations of an optical receiver system

Features:

Large collection area (16" diameter) telescope

Small area detector (200 μ m) Full beam overlap at ~ 0.5 km

Considerations
Receiver FOV
Biaxial or co-axial
F# of telescope
Near and far field signals
Alignment stability



Atmospheric Tests of Phototransistors

- Atmospheric tests conducted at NCAR (June 5-9, 2006) to test the sensitivity of small area (200 μ) phototransistors.
- High sensitivity Aerosol Scanning Lidar (REAL) system of NCAR's at Boulder, CO was used for atmospheric testing of new detectors.
- REAL operates in the 1.5 micron region and the phototransistor has reduced sensitivity in this region

Date: June 8, 2006

Time: 16:22:09 GMT

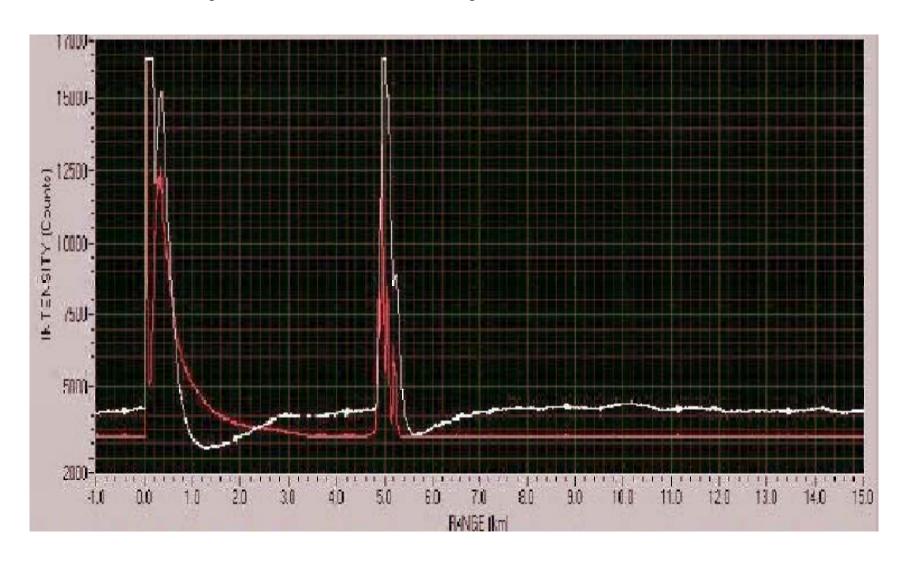
File: pic6 Elevation: 90 Azimuth: 310 Detector: LPE A1-b10 Raw

Bias Voltage: 3.5VTemperature: 20oCTIA Gain: 1000Ω Amp. Gain: 30dB File: REAL.20060608_153255.bscan

Comments: Looking at thin cirrus cloud at 5 km.

The cloud gets thinner. Even the boundary layer, at about 3 km, can be seen with the

phototransistor.



Date: June 8, 2006 Time: 16:40:44 GMT

File: pic10 Elevation: 90 Azimuth: 310 Detector: LPE A1-b10

Bias Voltage: 3.5V Temperature: 10oC TIA Gain: 1000Ω

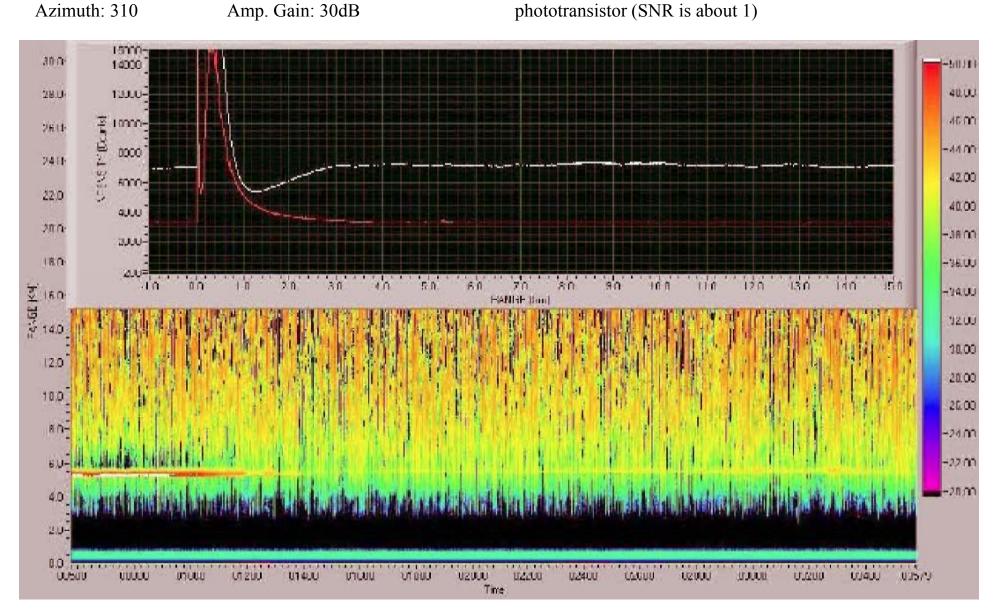
Raw File: REAL.20060608 162626.bscan

Comments: Cooling the phototransistor to 10°C

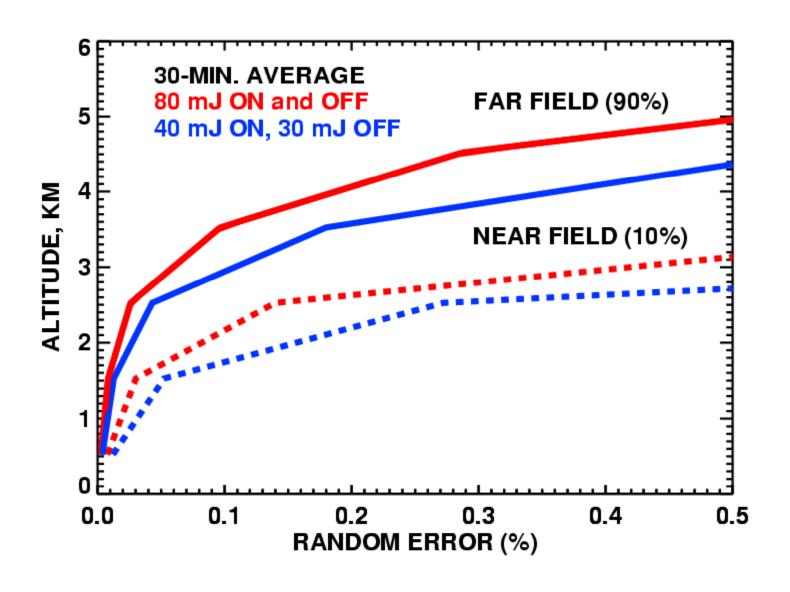
Although the feature becomes very weak (20% above bkgd.)

But still can be detected easily with the

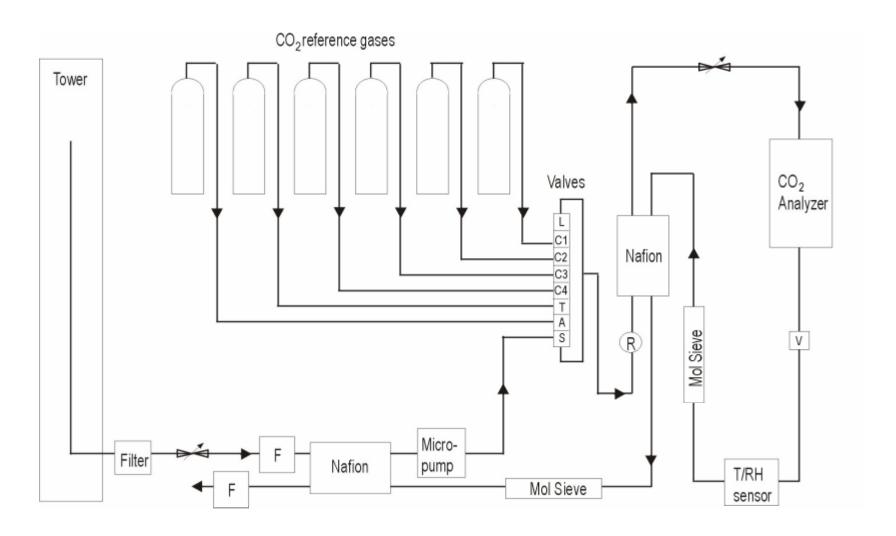
phototransistor (SNR is about 1)



Projected performance of two channel system

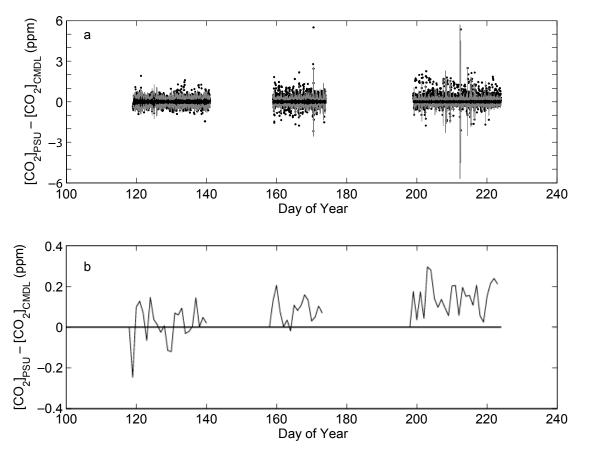


The PSU-NCAR CO₂ sensor design



For more information, see www.amerifluxco2.psu.edu

Performance Testing



Difference of daily averages

Difference between the PSU system and WLEF 76m CO₂ measurements in a test from April-August 2004. [Miles/Richardson/Uliasz, in prep.]

High Accuracy Analysis of R30 Intensity

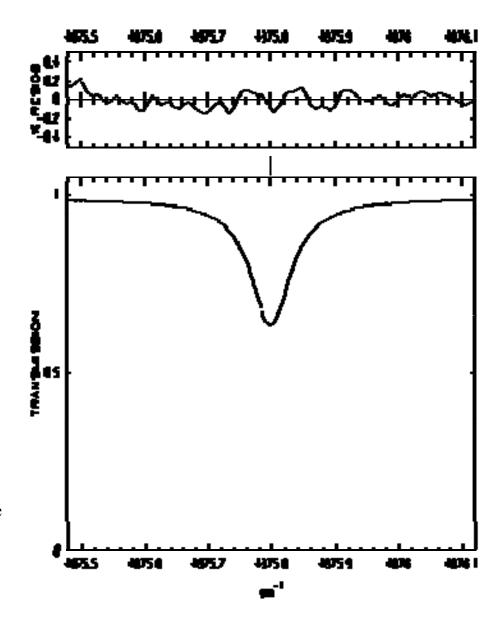
Progress to date:

High accuracy determination of Voigt line shape parameters

- Line position
- Line strength
- Pressure-dependent shift and width (self-induced)

Bottom Panel: Short scan of the CO_2 spectrum centered on R(30) [20013 \leftarrow 00001] with the experimental conditions as follows: sample pressure of 269.03 Torr, optical path of 0.347 m and sample temperature of 297.04K.

Top Panel: Residuals from simulating the experimental spectrum. Note that the residuals are on the order of 0.2%, a factor of 10 smaller than required, although there are some small systematic residuals around the line center characteristic of non-Voigt line shape behavior.



New R(30) Spectral Parameters

Line	Position	Strength	o-c(%)	σ(%)	b0	∆b0(%)	d0	∆d0(%)
R30	4875.74873	1.50E-22	0.06	0.34	0.0897	2.2	-7.16	1.8

Computed and observed minus computed, (o-c)%, line strengths of $^{12}C^{16}O_2$ in cm $^{-1}$ /(molecule cm $^{-2}$) at 296 K and standard mixture (0.9842 $^{12}C^{16}O_2$). σ (%) given relative to the strength value of Regalia-Jarlot et al. (2006).

Measured self-broadened $^{12}C^{16}O_2$ width coefficients b^0 with experimental width uncertainties, Δb^0 (in percent). Self-broadened half widths (HWHM) in units of cm⁻¹/atm at 296 K are in natural abundance (0.9842 $^{12}C^{16}O_2$).

Measured self-induced $^{12}C^{16}O_2$ shift coefficients, d⁰ with shift coefficient uncertainties, Δd^0 (in percent). Self-induced pressure shift coefficients are in units of cm⁻¹/atm x 10³ at 296 K for natural abundance (0.9842 $^{12}C^{16}O_2$).

Data from R. A. Toth, L. R. Brown, C. E. Miller V. Malathy Devi, and D. C. Benner, *J. Mol. Spectrosc.*, in press (2006)

Conclusions

- System will operate at a temperature insensitive CO₂ line (2050.967 nm) with side-line tuning and off-set locking.
- Demonstrated an order of magnitude improvement in laser line locking needed for high precision measurements, side-line operation, and simultaneously double pulsing and line locking.
- Detector testing of phototransistor has demonstrated sensitivity to aerosol features over long distances in the atmosphere and resolve features ~ 100m.
- Optical systems that collect light onto small area detectors work well.
- Receiver optical designs are being optimized and data acquisition systems developed.
- CO₂ line parameter characterization in progress
- In situ sensor calibration in progress for validation of DIAL CO₂ system

Lidar parameters of a direct detection CO₂ DIAL with Phototransistor and large 16" telescope

Baseline DIAL system parameters

Pulse energy = 80 mJ single pulsed; 40/30 mJ double pulsed

Pulse width = 180 ns

Pulse repetition rate = 5 Hz

Spectrum = single frequency

On-line wavelength = 2050.967 nm

off-line wavelength = 2050.940 nm

Beam quality < 1.3 time diffraction limit

Long term (one hour) wavelength stability < 2 MHz

Wavelength accuracy < 0.5 MHz

Detector AlGaAsSb/InGaAsSb phototransistor

Quantum efficiency = 70%

Optical efficiency 60%

NEP = 5.0E-14 W/sqrt(Hz)

Telescope aperture = 16 inches

A receiver system with two detectors to collect near- and far- field signals is planned—the collected signal is split in the ratio 9:1 between the near and far field detector channels

Error Budget

Random error from signal shot noise and detector noise		0.5%
Systematic errors: CO2 line temperature sensitivity		0.2%
CO2 cross-section error (residua	al)	0.2%
Detector influence (uncompensa	ted)	0.2%
Atmospheric number density erro	or	0.05%
H2O estimated dry air error		0.05%
Laser spectral error Total error (rms)	0.61%	0.05%